

Multiradionuclide Evidence for a Solar Energetic Particle Event at 14,300 BP

Solar Energetic Particle (SEP) events, resulting from intense solar eruptions, can cause short-term increases in the production of cosmogenic radionuclides such as ^{10}Be , ^{14}C , and ^{36}Cl . This study investigates whether such events occurred around 14,300 and 14,700 BP, intervals where pronounced spikes were previously observed in ^{10}Be and ^{14}C records.

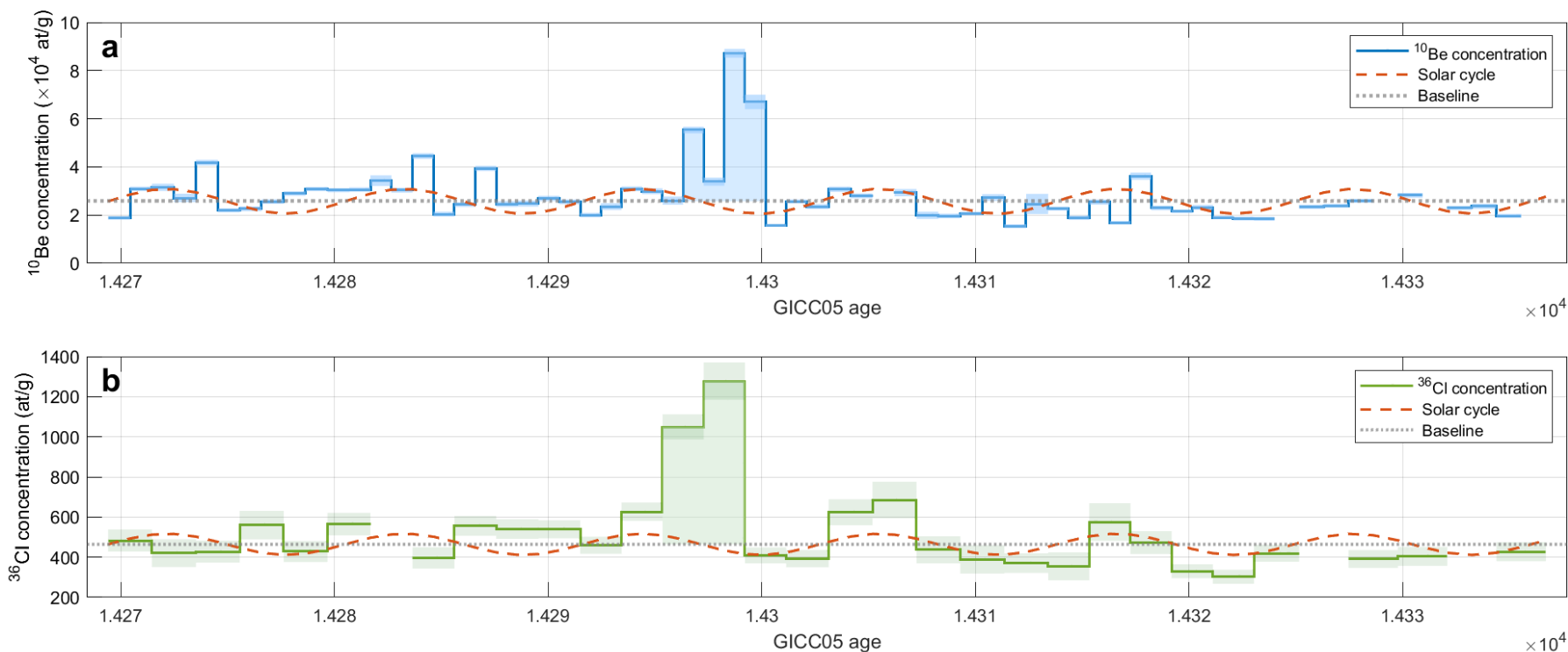


Fig. 1. NGRIP2 ^{10}Be and ^{36}Cl concentrations plotted against GICC05 age for the time interval 14,269.4–14,336.7 BP. (a) and (b) show ^{10}Be and ^{36}Cl concentrations, respectively, with associated uncertainties represented by shaded bars. The peak region is also shaded. Baselines and a reconstructed solar cycle are overlain.

To explore this further, high-resolution ^{10}Be and ^{36}Cl data from the North Greenland Ice Core Project was analysed. The results reveal one of the largest production peaks identified to date in both ^{10}Be and ^{36}Cl at 14,300 BP (Fig. 1). This peak coincides with a sharp rise in modelled ^{14}C production. This consistent signal across three radionuclide records provides strong evidence for an extreme SEP event. Variations in the ^{10}Be and ^{36}Cl concentrations exhibit an ~ 11 -year sinusoidal periodicity, consistent with modulation by the solar cycle. This places the event within a period of high solar activity, aligning with modern observations that they are most common near solar maxima. The event's $^{36}\text{Cl}/^{10}\text{Be}$ ratio suggests a hard energy spectrum dominated by very high-energy particles. Fluence estimates indicate the event was of an exceptional magnitude (F_{30} : $8.85 (\pm 0.86) \times 10^{10}$ particles/ cm^2), far exceeding any events observed during the instrumental era. These offer new insights into the timing, magnitude, and characteristics of extreme solar storms.

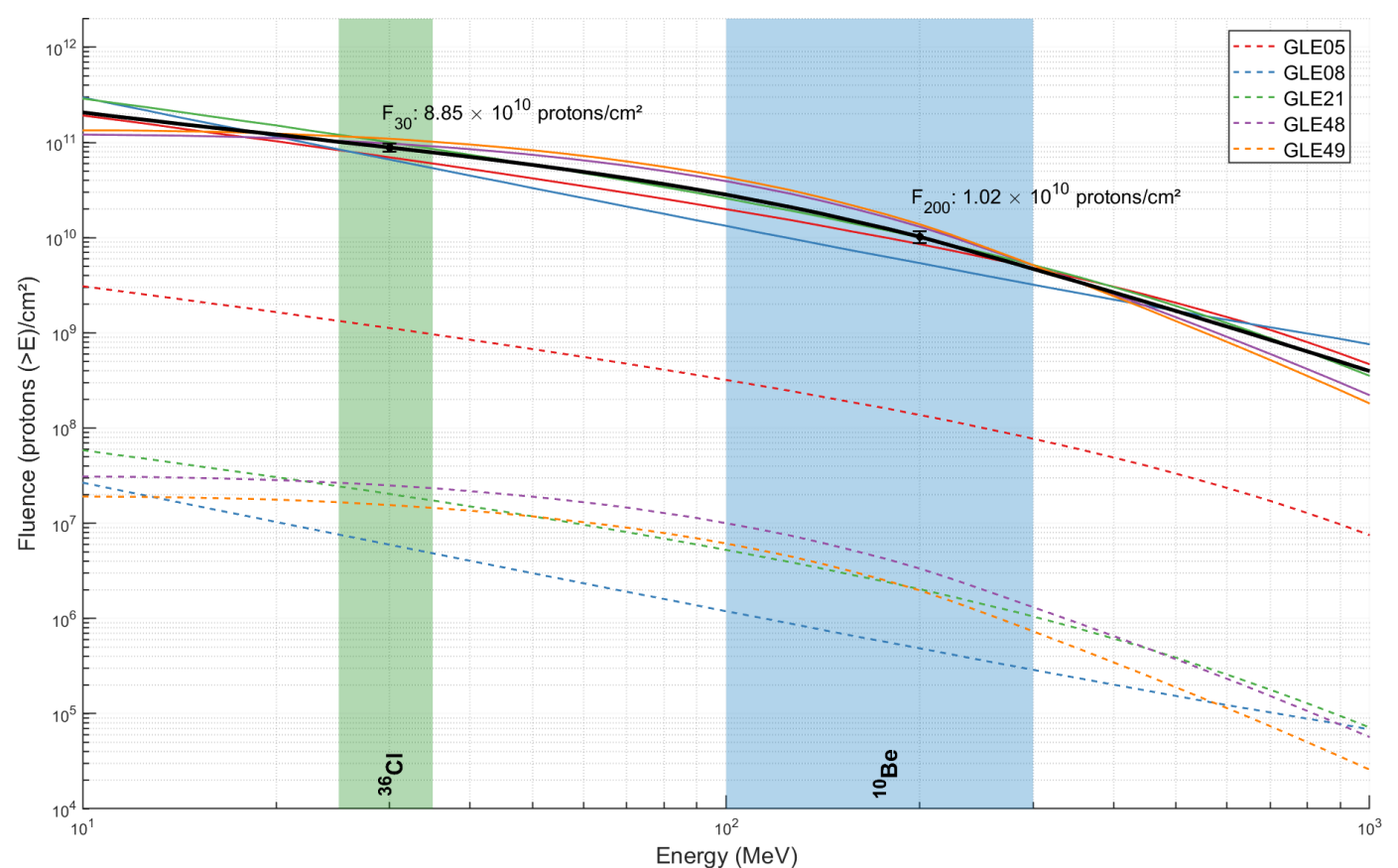


Fig. 2. Estimated fluence spectrum of the 14,300 BP event. Dashed lines show original GLE spectra that fall within the uncertainty range of the $^{36}\text{Cl}/^{10}\text{Be}$ ratio. Solid lines represent scaled versions of these spectra. The black solid curve indicates the average scaled spectrum (with F_{30} and F_{200} displayed), interpreted as the best estimate for the 14,300 BP event. Adapted from Koldobskiy et al., 2021; F. Mekhaldi (personal communications, April 25, 2025).

In contrast, the interval around 14,700 BP exhibits greater variability and lacks an unambiguous excursion. A significant correlation with $\delta^{18}\text{O}$ implies that these fluctuations are likely driven by climatic variability than solar activity.

Reference: Koldobskiy, S., Raukunen, O., Vainio, R., Kovaltsov, G.A., & Usoskin, I. (2021). New reconstruction of event-integrated spectra (spectral fluences) for major solar energetic particle events. *Astronomy & Astrophysics*, 647, A132.